

Physics potential of the ESS ν SB facility

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EUROPEAN SPALLATION SOURCE
NEUTRINO
SUPER BEAM

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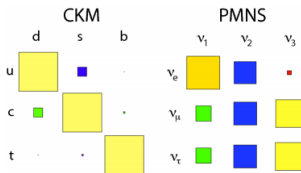
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Introduction

In the **SM neutrinos** \rightarrow **massless**
Neutrino oscillations \rightarrow **Physics BSM**

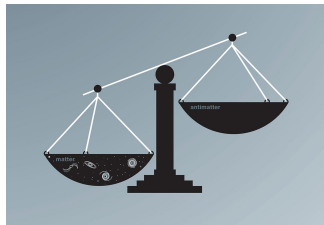
Flavour puzzle

- Why are fermion masses so different in the SM?
- Mixing for quarks very different from leptons



Baryon asymmetry

- Not enough CP violation in quark sector
- ν 's could be the answer through leptogenesis



What do we know?

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \text{diag}(1, e^{i\alpha_{21}}, e^{i\alpha_{31}})$$

Best fit values (1σ)

- $\Delta m_{21}^2 = 7.50_{-0.17}^{+0.19} \cdot 10^{-5} \text{eV}^2$
- $\sin^2 \theta_{12} = 0.306_{-0.012}^{+0.012}$
- $|\Delta m_{31}^2| = 2.524_{-0.040}^{+0.039} \cdot 10^{-3} \text{eV}^2$
- $\sin^2 \theta_{23} = 0.441_{-0.021}^{+0.027}$
- $\sin^2 \theta_{13} = 0.02166_{-0.0075}^{+0.00075}$

I. Esteban, M. Gonzalez-Garcia, M. Maltoni, I. Martinez-Soler, T. Schwetz hep-ph/1611.01514

Still unknown

- $\delta \rightarrow$ Leptonic CP violation
- $Sign(\Delta m_{31}^2)$
- Octant of θ_{23}
- ν 's Dirac or Majorana?
Not testable with ν oscillations

The interference term

In matter

A. Cervera *et al.* hep-ph/0002108

$$P_{\mu \rightarrow e}^{\pm} = s_{23}^2 \sin^2 2\theta_{13} \left(\frac{\Delta_{atm}}{\tilde{B}_{\mp}} \right)^2 \sin^2 \left(\frac{\tilde{B}_{\mp} L}{2} \right) \text{ Atmospheric}$$

$$+ c_{23}^2 \sin^2 2\theta_{12} \left(\frac{\Delta_{sol}}{A} \right)^2 \sin^2 \left(\frac{AL}{2} \right) \text{ Solar}$$

$$\text{Interference} + J \frac{\Delta_{sol}}{A} \frac{\Delta_{atm}}{\tilde{B}_{\mp}} \sin \left(\frac{AL}{2} \right) \sin \left(\frac{\tilde{B}_{\mp} L}{2} \right) \cos \left(\pm \delta - \frac{\Delta_{atm} L}{2} \right)$$

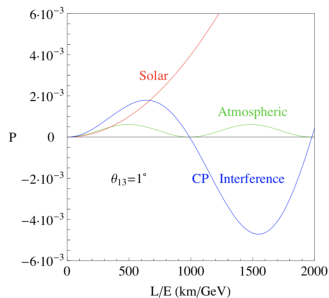
$$J = \cos \theta_{13} \sin 2\theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \quad \Delta_{atm} = \frac{\Delta m_{31}^2}{2E} \quad \Delta_{sol} = \frac{\Delta m_{12}^2}{2E}$$

$$A = \sqrt{2} G_F n_e \quad \tilde{B}_{\mp} = |A \mp \Delta_{atm}|$$

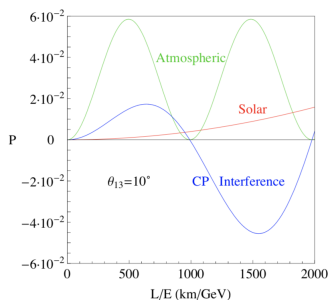
Impact of large θ_{13}

P. Coloma, E. Fernandez-Martinez hep-ph/1110.4583

θ_{13} small



θ_{13} large



First maximum:

- Large statistics
- Systematic error as bottleneck

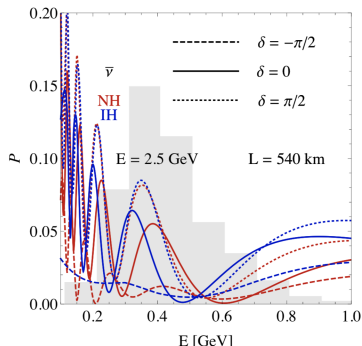
Second maximum:

- Systematic error not a problem
- Statistics as bottleneck

The ESS ν SB

E. Baussan *et al.* hep-ex/1309.7022

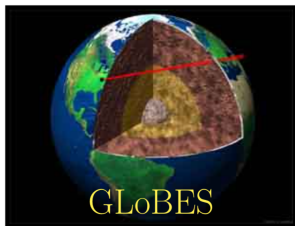
- Modification of the ESS linac to produce neutrinos 14Hz \rightarrow 28Hz
- 5 MW at 2.5GeV proton beam
- MEMPHYS-like WC detector
 - 500 kt fiducial volume
 - Best locations at **540** and 360 km



Simulation framework

P. Huber *et al.* hep-ph/0701187

- Implemented in GLoBES
- Simulation with a ND
- 2.5 GeV proton beam



P. Coloma *et al.* hep-ph/1209.5973

Systematic errors

Systematics	SB		
	Opt.	Def.	Cons.
Fiducial volume ND	0.2%	0.5%	1%
Fiducial volume FD (incl. near-far extrap.)	1%	2.5%	5%
Flux error signal ν	5%	7.5%	10%
Flux error background ν	10%	15%	20%
Flux error signal $\bar{\nu}$	10%	15%	20%
Flux error background $\bar{\nu}$	20%	30%	40%
Background uncertainty	5%	7.5%	10%
Cross secs \times eff. QE †	10%	15%	20%
Cross secs \times eff. RES †	10%	15%	20%
Cross secs \times eff. DIS †	5%	7.5%	10%
Effec. ratio ν_e/ν_μ QE *	3.5%	11%	—
Effec. ratio ν_e/ν_μ RES *	2.7%	5.4%	—
Effec. ratio ν_e/ν_μ DIS *	2.5%	5.1%	—
Matter density	1%	2%	5%

Sensitivity to mass hierarchy

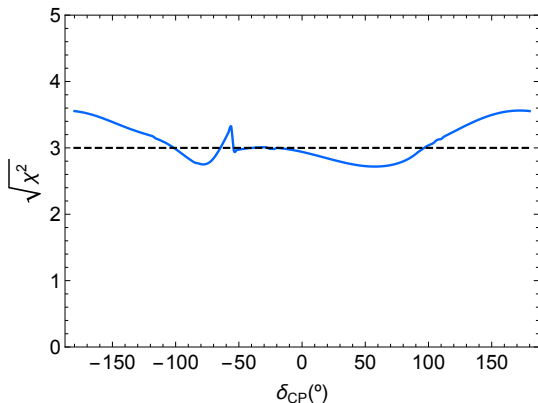
Matter effects relevant when

- High energy ν
 $E_\nu \sim 6 \text{ GeV}$

ESS ν SB

- $L \sim 500 \text{ km}$
- $E_\nu \sim \mathcal{O}(300 \text{ MeV})$

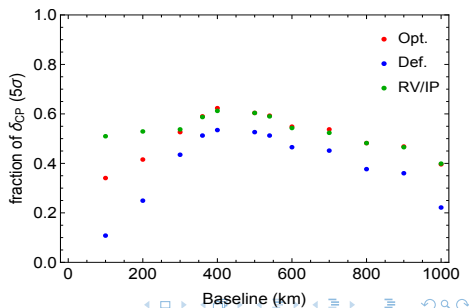
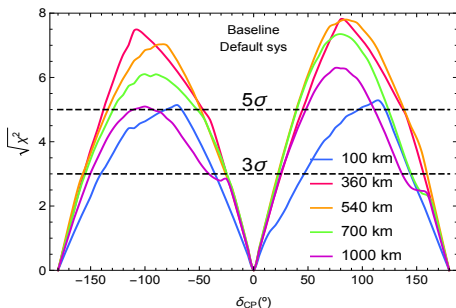
Sensitivity around 3σ



Sensitivity to CP violation

Comparison between

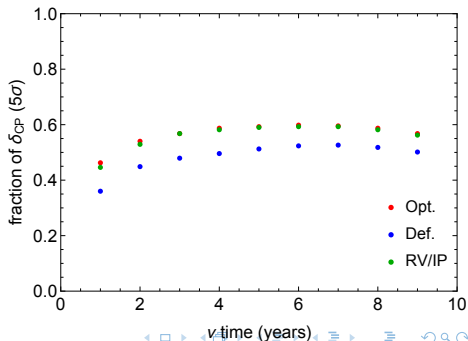
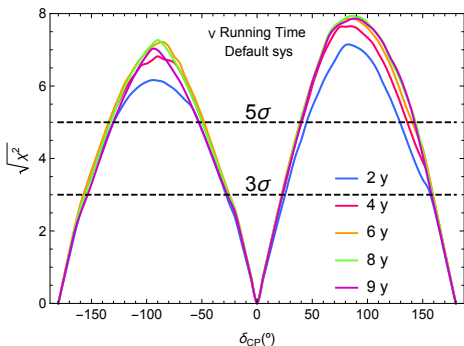
- “Default” systematics
- “Optimistic” systematics
- $T2HK$ $\mathcal{O}(3\%)$ systematics with 374 kt (reduced volume and improved photocoverage) [K. Abe et al. 1805.04163](#)



Sensitivity to CP violation

Impact of ν mode running time

- Up to now studies with 2+8 years
- Preference for 5+5 years



Sensitivity to CP violation

Impact of each systematic error

For different values of the systematics

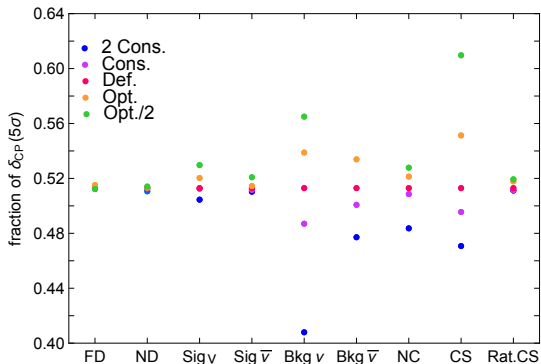
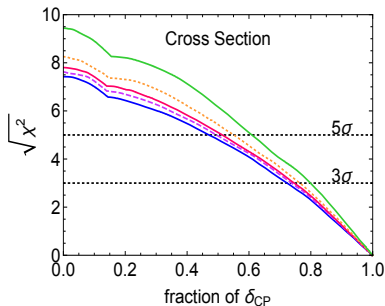
- $2 \times$ Cons.
- Cons.
- Def.
- Opt.
- Opt./2

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Effec. ratio ν_e/ν_μ RES [*]	2.7%	5.4%	–
Effec. ratio ν_e/ν_μ DIS [*]	2.5%	5.1%	–
Matter density	1%	2%	5%

Sensitivity to CP violation

Most important systematics

- ν background
- Cross section uncertainties



Conclusions

- **ESS ν SB**

Complementarity with *DUNE*. **Lower energy ν** .

- Mass hierarchy

It could exclude the wrong hierarchy at $\sim 3\sigma$

- ν mode running time

Better a symmetric **5+5 setup**

- Baseline

- $L < 300\text{km}$ systematics relevant \rightarrow RV/IP better as $\mathcal{O}(3\%)$ sys.

- $L_{\text{optimal}} \sim 400\text{km}$ for **5 + 5** yrs running

- Impact of systematic errors

ν **background** and **cross section**

$L = 540 \text{ km}$ and **5 + 5 yrs**

CP violation could be discovered above **5 σ** for a **51.2 – 59.2%** of values of δ depending on the systematics